

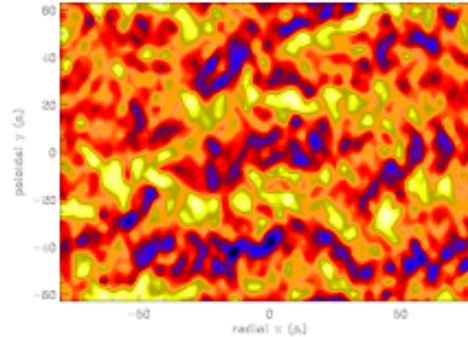
# Overview of Gyrofluid Equations and their Numerical Solution

**P.B. Snyder<sup>1</sup> and G.W. Hammett<sup>2</sup>**

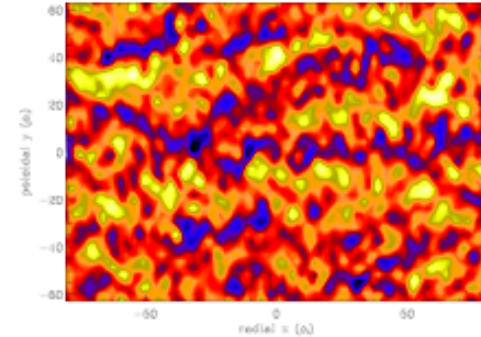
<sup>1</sup>*General Atomics, San Diego, USA*

<sup>2</sup>*Princeton Plasma Physics Laboratory, Princeton, USA*

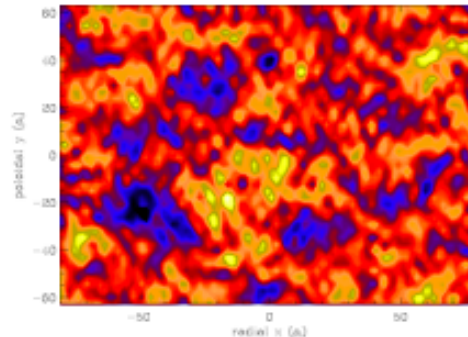
$\beta = 0\%$



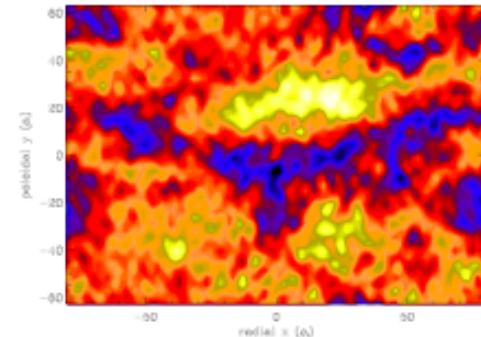
$\beta = .4\%$



$\beta = .8\%$



$\beta = 1\%$



BOUT++ Workshop  
16 September 2011

# Definition of Gyrofluid



- **“Landau Fluid” equations are fluid moment equations which are truncated with closure terms incorporating kinetic effects**
  - Landau damping (linear), Toroidal drift resonance...
- **“Gyro-Landau Fluid” or “Gyrofluid” equations are fluid moments of the 5D gyrokinetic equation, closed so as to maintain FLR and kinetic effects**
  - Landau damping, linear and nonlinear FLR, toroidal drifts and drift resonance, trapping
  - Able to accurately reproduce linear GK physics, and provide reasonable agreement with nonlinear GK simulations, while being much more efficient
  - Eg: GRYFFIN (Beer, Dorland, Hammett, Snyder) and Waltz GLF simulation codes, GLF23 and TGLF linear models

# Vlasov, Boltzmann, Liouville Eq:

Particle Distribution

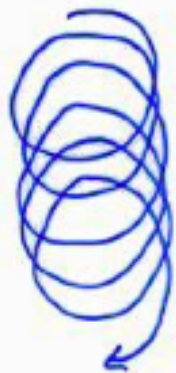
$$\frac{\partial f(\underline{x}, \underline{v}, t)}{\partial t} + \underline{v} \cdot \frac{\partial f}{\partial \underline{x}} + \frac{q}{m} \left( \underline{E} + \frac{\underline{v} \times \underline{B}}{c} \right) \cdot \frac{\partial f}{\partial \underline{v}} = C(f)$$

Nonlinear,  $\underline{E} \& \underline{B}$  depend on  $f$  through Maxwell's Eqs.

## Nonlinear Gyrokinetic Eq. 1982-88

(Frieman & Chen, W.W. Lee, Dubin, Krommes, Hahn, Brizard...)

linear gyrokinetics  
1960's & 70's.



=



+



$\underline{v}_{drift}$

$$\frac{c}{B^2} \langle \underline{E} \rangle \times \underline{B} + \dots$$

Possible to eliminate fast gyrofrequency  $\Phi$   
time scales & retain nonlinear dynamics

$$\& k_{\perp} \rho_i \sim 1$$



# Comparison of Gyrofluid and Traditional MHD-like 2-fluid Equations

- **GF: Moments of 5D GK eqn rather than 6D K eqn**
  - Fast timescales and short scale lengths eliminated before moments are taken  $\frac{\omega}{\Omega_i} \sim \frac{k_{\perp} v_{th}}{\Omega_i} \sim \frac{e\phi}{T} \sim \frac{\delta B}{B} \sim \frac{F_1}{F_0} \sim \frac{\rho_i}{L} \sim \varepsilon \ll 1, \quad k_{\perp} \rho_i \sim 1,$
- **GF: Moments taken in gyro-center space**
  - Gyro-viscous cancellation natural, algebra easier
- **GF: Closures derived by matching to linear kinetic response, rather than high collisionality**
  - Can accurately reproduce kinetic physics in both low and high collisionality limits
- **GF: FLR and closure terms take a form which can be efficiently evaluated in k-space**
  - Much more challenging to evaluate in x-space

# Deriving Gyrofluid Equations

- Start with GK eqn:

$$\frac{\omega}{\Omega_i} \sim \frac{k_{\parallel} v_{ti}}{\Omega_i} \sim \frac{e\phi}{T} \sim \frac{\delta B}{B} \sim \frac{F_1}{F_0} \sim \frac{\rho_i}{L} \sim \varepsilon \ll 1, \quad k_{\perp} \rho_i \sim 1, \quad \frac{\partial F}{\partial t} + (v_{\parallel} \tilde{\mathbf{b}} + \mathbf{v}_E + \mathbf{v}_d) \cdot \nabla F + \left[ \frac{e}{m} \tilde{E}_{\parallel} - \mu \tilde{\mathbf{b}} \cdot \nabla B + v_{\parallel} (\tilde{\mathbf{b}} \cdot \nabla \tilde{\mathbf{b}}) \cdot \mathbf{v}_E \right] \frac{\partial F}{\partial v_{\parallel}} = C(F).$$

- Put it in conservative form:

$$\begin{aligned} \frac{\partial \tilde{f} B}{\partial t} &+ B \nabla_{\parallel} \tilde{f} v_{\parallel} + \mathbf{v}_{\phi} \cdot \nabla [(F_0 + \tilde{f}) B J_0] - \mathbf{v}_{A_{\parallel}} \cdot \nabla [(F_0 + \tilde{f}) B \frac{v_{\parallel}}{c} J_0] \\ &+ 2F_0 B J_0 i\omega_d \frac{e\phi}{T} - F_0 B \frac{v_{\parallel}}{c} J_0 i\omega_d \frac{eA_{\parallel}}{T} + F_0 B J_1 \frac{\alpha}{2} i\omega_d \frac{e\phi}{T} \\ &- F_0 B \frac{v_{\parallel}}{c} J_1 \frac{\alpha}{2} i\omega_d \frac{eA_{\parallel}}{T} + \frac{i\omega_d}{v_{\parallel}^2} [\tilde{f} B (v_{\parallel}^2 + \mu B)] - \frac{e}{mc} \frac{\partial F_0}{\partial v_{\parallel}} B J_0 \frac{\partial A_{\parallel}}{\partial t} \\ &- \frac{e}{m} \nabla_{\parallel} \left( \frac{\partial F_0}{\partial v_{\parallel}} B J_0 \phi \right) + \frac{e}{m} J_0 \phi \frac{\partial F_0}{\partial v_{\parallel}} B \left( 1 - \frac{\mu B}{v_{\parallel}^2} \right) \nabla_{\parallel} \ln B \\ &+ \frac{e}{m} \frac{\partial F_0}{\partial v_{\parallel}} J_0 A_{\parallel} J_0 \phi (\tilde{\mathbf{b}} \times \nabla A_{\parallel}) \cdot \nabla \phi - \mu B^2 \frac{\partial \tilde{f}}{\partial v_{\parallel}} \nabla_{\parallel} \ln B \\ &- \frac{\partial F_0}{\partial v_{\parallel}} B J_0 \frac{\mu B}{c} i\omega_d \frac{eA_{\parallel}}{T} - \frac{\partial}{\partial v_{\parallel}} (F_0 B J_0 v_{\parallel}) i\omega_d \frac{e\phi}{T} = 0. \end{aligned}$$

- Take moments:

Velocity space moments are often defined in terms of the total distribution function  $F$ . Here we again separate  $F$  into equilibrium and fluctuating components  $F = F_0 + \tilde{f}$ .<sup>6</sup> Velocity space moments of

$$F_0 = F_M = \frac{n_0}{(2\pi v_{ti}^2)^{3/2}} e^{-v_{\parallel}^2/2v_{ti}^2 - \mu B/v_{ti}^2}$$

are all well defined. We define the following moments of the fluctuating distribution:

$$\begin{aligned} \tilde{n} &= \int \tilde{f} d^3v & n_0 \tilde{u}_{\parallel} &= \int \tilde{f} v_{\parallel} d^3v \\ \tilde{p}_{\parallel} &= m \int \tilde{f} v_{\parallel}^2 d^3v & \tilde{p}_{\perp} &= m \int \tilde{f} B \mu d^3v \\ \tilde{q}_{\parallel} &= -3m v_{ti}^2 n_0 \tilde{u}_{\parallel} + m \int \tilde{f} v_{\parallel}^3 d^3v & \tilde{q}_{\perp} &= -m v_{ti}^2 n_0 \tilde{u}_{\parallel} + m \int \tilde{f} B \mu v_{\parallel} d^3v \\ \tilde{r}_{\parallel,\perp} &= m \int \tilde{f} v_{\parallel}^4 d^3v & \tilde{r}_{\parallel,\perp} &= m \int \tilde{f} B \mu v_{\parallel}^2 d^3v \\ \tilde{r}_{\perp,\perp} &= m \int \tilde{f} B^2 \mu^2 d^3v & \tilde{s}_{\perp,\perp} &= -2m v_{ti}^4 n_0 \tilde{u}_{\parallel} + m \int \tilde{f} B^2 \mu^2 v_{\parallel} d^3v \\ \tilde{s}_{\parallel,\perp} &= -15m v_{ti}^4 n_0 \tilde{u}_{\parallel} + m \int \tilde{f} v_{\parallel}^5 d^3v & \tilde{s}_{\parallel,\perp} &= -3m v_{ti}^4 n_0 \tilde{u}_{\parallel} + m \int \tilde{f} B \mu v_{\parallel}^3 d^3v, \end{aligned}$$

Closure of high moments (3+1 or 4+2) preserves particle, momentum and energy conservation



# Deriving Closure Terms

- Closures needed for FLR, parallel, toroidal drift and mirror terms
- Landau damping:

As an illustration, consider the one dimensional kinetic equation

$$\frac{\partial f}{\partial t} + v_{\parallel} \frac{\partial f}{\partial z} = \delta(t) f_0(z, v), \quad (3.74)$$

Nothing is really damped in Landau damping.

Phase mixing moves fluctuations to fine scales in  $v$  space.

Once at small scales we assume they are damped by collisions. Good assumption for turbulence, not so good for special cases like plasma echo

where  $f_0$  provides the initial condition. The solution to this simple equation  $f(z, v, t) = f_0(z - vt, v)H(t)$ , provides Green's function which can be used to solve more general problems with additional source terms, such as the electric field  $-(e/m)E_{\parallel} \frac{\partial F_M}{\partial v}$ . Consider an initial condition with a small single harmonic perturbation  $f_0 = (n_0 + n_1 e^{ikz})F_M(v)$ . The general solution is just  $(n_0 + n_1 e^{ik(z-vt)})$ , which simply oscillates in time at  $\omega = kv$  and does not damp. However, upon taking velocity space moments, the velocity integration introduces mixing of the phases as follows:

$$n(z, t) = \int f dv = n_0 + n_1 \frac{e^{ikz}}{\sqrt{2\pi v_t^2}} \int dv e^{-ikvt} e^{-v^2/(2v_t^2)}. \quad (3.75)$$

The perturbed density  $n_1 = n_{1(t=0)} e^{-k^2 v_t^2 t^2/2}$  decays with a Gaussian time dependence. This decay due to linear Landau damping is not captured by a simple fluid model with a finite number of moments, and hence it must be accounted for in the fluid closure if it is to be included in a fluid model.

# Deriving Closure Terms: Practice

- Kinetic linear response:

$$n_{1s} = -\frac{in_0}{k_z T_{\parallel 0s}} e_s E_{\parallel} \mathcal{R}(\zeta_s) + \frac{B_1 n_0}{B_0} \left[ 1 - \frac{T_{\perp 0s}}{T_{\parallel 0s}} \mathcal{R}(\zeta_s) \right]$$

where  $\zeta_s = \omega/\sqrt{2}|k_z|v_{t\parallel s}$  is the normalized frequency, and  $\mathcal{R}(\zeta_s) = 1 + \zeta_s Z(\zeta_s)$

- GF closure form:
- D' s determined by matching K response in small and large z limit

$$r_{\parallel, \perp s} = 3v_{t\parallel s}^2 (2p_{\parallel s} - T_{\parallel 0s} n) + c_{\parallel} n_0 v_{t\parallel s}^2 T_{\parallel s} - \sqrt{2} D_{\parallel} v_{t\parallel s} \frac{ik_{\parallel} q_{\perp s}}{|k_{\parallel}|}$$

$$r_{\parallel, \perp s} = v_{t\perp s}^2 p_{\parallel s} + v_{t\parallel s}^2 p_{\perp s} - v_{t\parallel s}^2 T_{\perp 0s} n - \sqrt{2} D_{\perp} v_{t\parallel s} \frac{ik_{\parallel} q_{\perp s}}{|k_{\parallel}|}$$

The density response is then:

$$n_{1s} = -\frac{in_0}{k_z T_{\parallel 0s}} e_s E_{\parallel} \mathcal{R}_4(\zeta_s) + \frac{B_1 n_0}{B_0} \left[ 1 - \frac{T_{\perp 0s}}{T_{\parallel 0s}} \mathcal{R}_4(\zeta_s) \right] \quad (\text{C.36})$$

where  $\mathcal{R}_4(\zeta_s)$  is a four-pole model of the electrostatic response function  $\mathcal{R}(\zeta_s)$ :

$$\mathcal{R}_4(\zeta_s) = \frac{4 - 2i\sqrt{\pi}\zeta_s + (8 - 3\pi)\zeta_s^2}{4 - 6i\sqrt{\pi}\zeta_s + (16 - 9\pi)\zeta_s^2 + 4i\sqrt{\pi}\zeta_s^3 + (6\pi - 16)\zeta_s^4} \quad (\text{C.37})$$

# Accurately reproduces kinetic response and linear growth rates

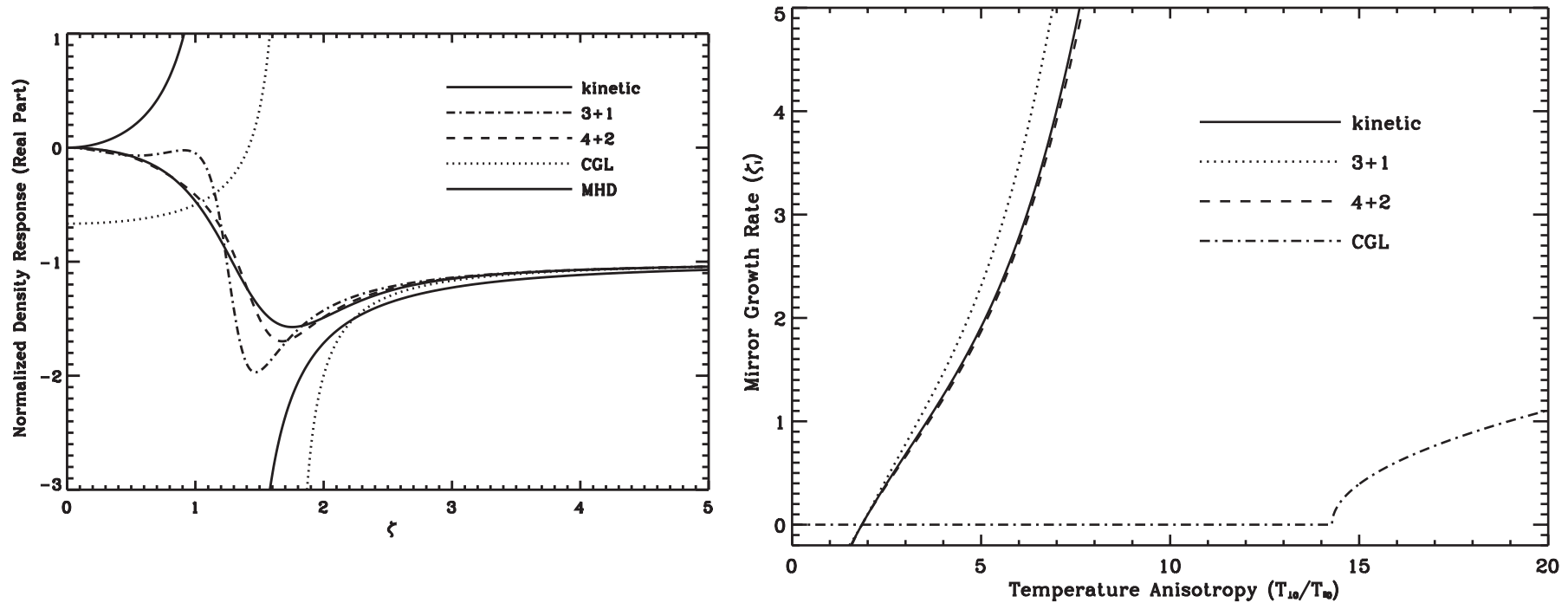


Figure C.3: Linear growth rate of the mirror instability ( $k^2 \gg k_{\parallel}^2$ ) as predicted by kinetic theory, 3+1 and 4+2 Landau MHD models, and CGL theory (ideal MHD cannot predict the mirror growth rate as it posits an isotropic pressure). The normalized growth rate ( $\zeta_i = \text{Im}(\omega) / \sqrt{2} |k_{\parallel}| v_{T_{\parallel i}}$ ) is plotted versus the temperature anisotropy ( $T_{\perp 0}/T_{\parallel 0}$ ) at constant  $\beta = \{(2/3)p_{\perp 0} + (1/3)p_{\parallel 0}\} / (B_0^2/8\pi)$ . The parameters chosen are  $Z = 1$ ,  $T_{\perp 0i} = T_{\perp 0e}$ ,  $T_{\parallel 0i} = T_{\parallel 0e}$ ,  $\beta = 1$  and  $\sqrt{m_i/m_e} = 40$ .



# Adding FLR, Toroidal Drift & Mirror Closures gives final EM Gyrofluid equations

## Toroidal Ion Gyrofluid Equations

$$\begin{aligned}
 \frac{dn}{dt} &+ \left[ \frac{1}{2} \tilde{\nabla}^2 \mathbf{v}_\psi \cdot \nabla T_i - \left[ \frac{1}{2} \tilde{\nabla}^2 \mathbf{v}_A \cdot \nabla q_i + B \tilde{\nabla}_i \frac{q_i}{R} - \left( 1 + \frac{\eta}{2} \tilde{\nabla}_i^2 \right) i\omega_s \Psi \right] \right. \\
 &+ \left. \left( 2 + \frac{1}{2} \tilde{\nabla}^2 \right) i\omega_d \Psi + i\omega_d (p_i + p_e) \right] = 0 \\
 \frac{dq_i}{dt} &+ \left[ \frac{1}{2} \tilde{\nabla}^2 \mathbf{v}_\psi \cdot \nabla q_i + B \tilde{\nabla}_i \frac{q_i}{R} + \frac{\partial A_i}{\partial t} + \tilde{\nabla}_i \Psi + \left( 1 + \eta_i + \frac{\eta}{2} \tilde{\nabla}^2 \right) i\omega_s A_i \right. \\
 &- \left. \left[ \frac{1}{2} \tilde{\nabla}^2 \mathbf{v}_A \cdot \nabla T_i + \left( p_i + \frac{1}{2} \tilde{\nabla}_i^2 \Psi \right) \nabla_i \ln B + i\omega_d (q_i + q_e + 4u_i) \right] \right] = 0 \\
 \frac{dq_e}{dt} &+ \left[ \frac{1}{2} \tilde{\nabla}^2 \mathbf{v}_\psi \cdot \nabla T_i + B \tilde{\nabla}_i \frac{q_i + 3u_i}{R} + 2(q_i + u_i) \nabla_i \ln B \right. \\
 &- \left. \left( 1 + \eta_i + \frac{\eta}{2} \tilde{\nabla}^2 \right) i\omega_s \Psi + \left( 4 + \frac{1}{2} \tilde{\nabla}^2 \right) i\omega_d \Psi + i\omega_d (7p_i + p_e - 4n) \right. \\
 &+ \left. 2|\omega_d|(\nu_- T_i + \nu_2 T_i) = -\frac{2}{3} \nu_{se}(p_i - p_e) \right] \\
 \frac{dp_i}{dt} &+ \left[ \frac{1}{2} \tilde{\nabla}^2 \mathbf{v}_\psi \cdot \nabla p_i + [\tilde{\nabla}_i^2 \mathbf{v}_\psi] \cdot \nabla T_i - \left[ \frac{1}{2} \tilde{\nabla}^2 \mathbf{v}_A \cdot \nabla (q_i + u_i) \right. \right. \\
 &+ \left. \left. B^2 \tilde{\nabla}_i \frac{q_i + u_i}{R^2} - \left[ 1 + \frac{1}{2} \tilde{\nabla}_i^2 + \eta_i \left( 1 + \frac{1}{2} \tilde{\nabla}^2 + \tilde{\nabla}_i^2 \right) \right] i\omega_s \Psi \right. \right. \\
 &+ \left. \left. \left( 3 + 2\tilde{\nabla}_i^2 + 3\tilde{\nabla}^2 \right) i\omega_d \Psi + \left( 3 + \frac{3}{2} \tilde{\nabla}_i^2 + \tilde{\nabla}_i^2 \right) i\omega_d \Psi \right. \right. \\
 &+ \left. \left. i\omega_d (5p_i + p_i - 3n) + 2|\omega_d|(\nu_3 T_i + \nu_4 T_i) = \frac{1}{3} \nu_{se}(p_i - p_e) \right] \right. \\
 \frac{dq_e}{dt} &+ \left[ 3\tilde{\nabla}_i T_i + \alpha_q \nabla_i T_i + \sqrt{2} D_\lambda |k_\parallel| q_i + i\omega_d (-3q_i - 3q_i + 6u_i) \right. \\
 &+ \left. 3\eta_i i\omega_s A_i + |\omega_d|(\nu_5 u_i + \nu_6 q_i + \nu_7 q_i) = -\nu_{se} q_i \right] \\
 \frac{dq_i}{dt} &+ \left[ \frac{1}{2} \tilde{\nabla}^2 \mathbf{v}_\psi \cdot \nabla u_i + [\tilde{\nabla}_i^2 \mathbf{v}_\psi] \cdot \nabla q_i - \left[ \tilde{\nabla}_i^2 - \frac{1}{2} \tilde{\nabla}_i^2 \right] \mathbf{v}_A \cdot \nabla T_i + \tilde{\nabla}_i T_i \right. \\
 &+ \left. \left[ \eta_i (1 + \tilde{\nabla}^2) + (1 + \eta_i) \frac{1}{2} \tilde{\nabla}^2 \right] i\omega_s A_i + \frac{1}{2} \tilde{\nabla}_i^2 \left( \frac{dA_i}{dt} + \tilde{\nabla}_i \Psi - i\omega_d A_i \right) \right. \\
 &+ \left. \sqrt{2} D_\lambda |k_\parallel| q_i + \left( p_i - p_i + \tilde{\nabla}_i^2 \Psi - \frac{1}{2} \tilde{\nabla}_i^2 \Psi \right) \nabla_i \ln B \right. \\
 &+ \left. i\omega_d (-q_i - q_i + u_i) + |\omega_d|(\nu_8 u_i + \nu_9 q_i + \nu_{10} q_i) = -\nu_{se} q_i \right]
 \end{aligned}$$

where  $\tilde{\nabla}_i = \nabla_i - \hat{\mathbf{b}} \times \nabla A_i \cdot \nabla$ ,  $\mathbf{v}_A = \hat{\mathbf{b}} \times \nabla A_i$ ,  $A_i = \Gamma_0^{1/2} A_i$ .

$$n_e = \bar{n}_i - (1 - \Gamma_0) \phi,$$

$$\nabla_\perp^2 A_i = -\frac{\tau \beta_e}{2} (\bar{u}_{i\perp} - u_{i\perp e}),$$

Can simplify for special cases,

eg low mass electrons:

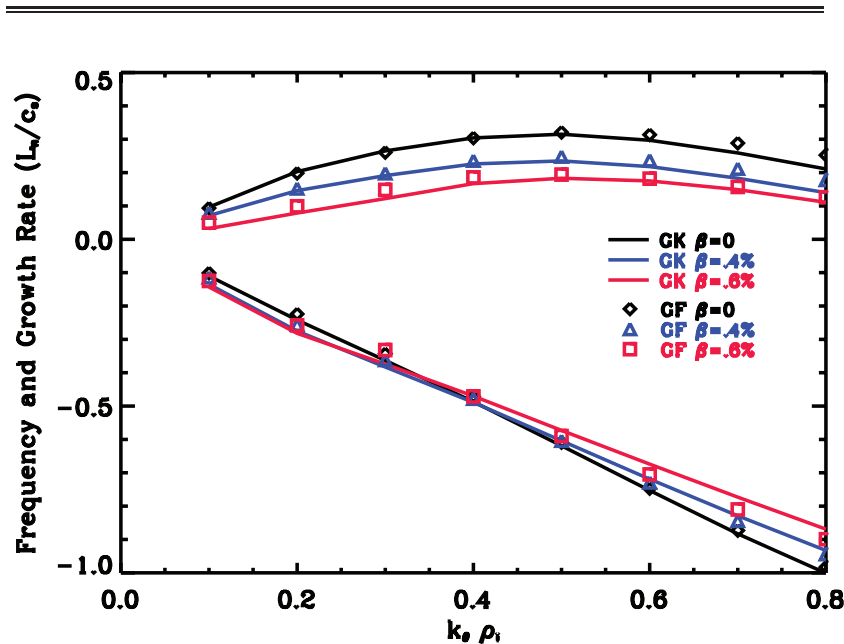
$$\frac{\partial n_e}{\partial t} + \mathbf{v}_E \cdot \nabla n_e + B \tilde{\nabla}_i \frac{u_{i\perp e}}{B} - i\omega_s \phi + 2i\omega_d \left( \phi - \frac{n_e}{\tau} - T_e \right) = 0,$$

$$\frac{\partial A_i}{\partial t} + \tilde{\nabla}_i \phi - \frac{1}{\tau} \tilde{\nabla}_i n_e - \frac{1}{\tau} i\omega_s A_i - \sqrt{\frac{\pi}{2\tau} \frac{m_e}{m_i}} |k_\parallel| u_{i\perp e} = \nu_{ei} \frac{m_e}{m_i} (u_{i\perp e} - u_{i\perp i}),$$

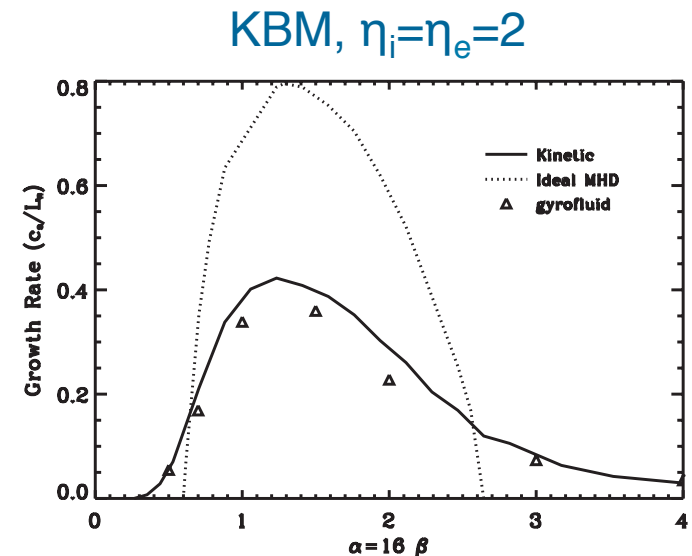
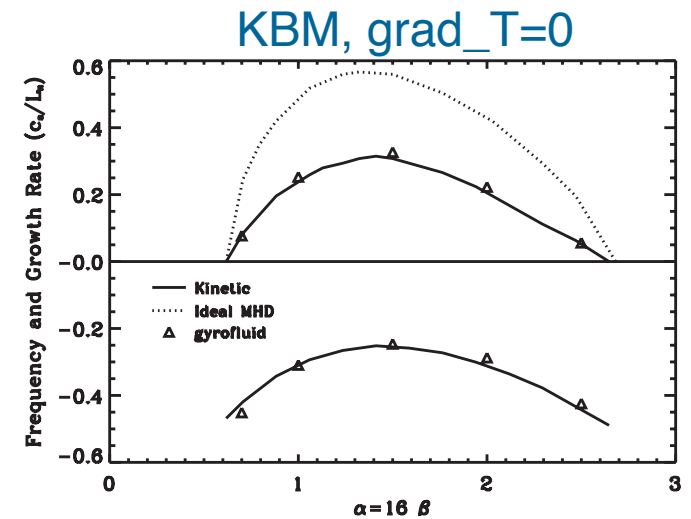
$$\tilde{\nabla}_i T_e = -\frac{\eta_e}{\tau} i\omega_s A_i.$$

# Adding FLR, Toroidal Drift & Mirror Closures allows accurate treatment of GK drift modes

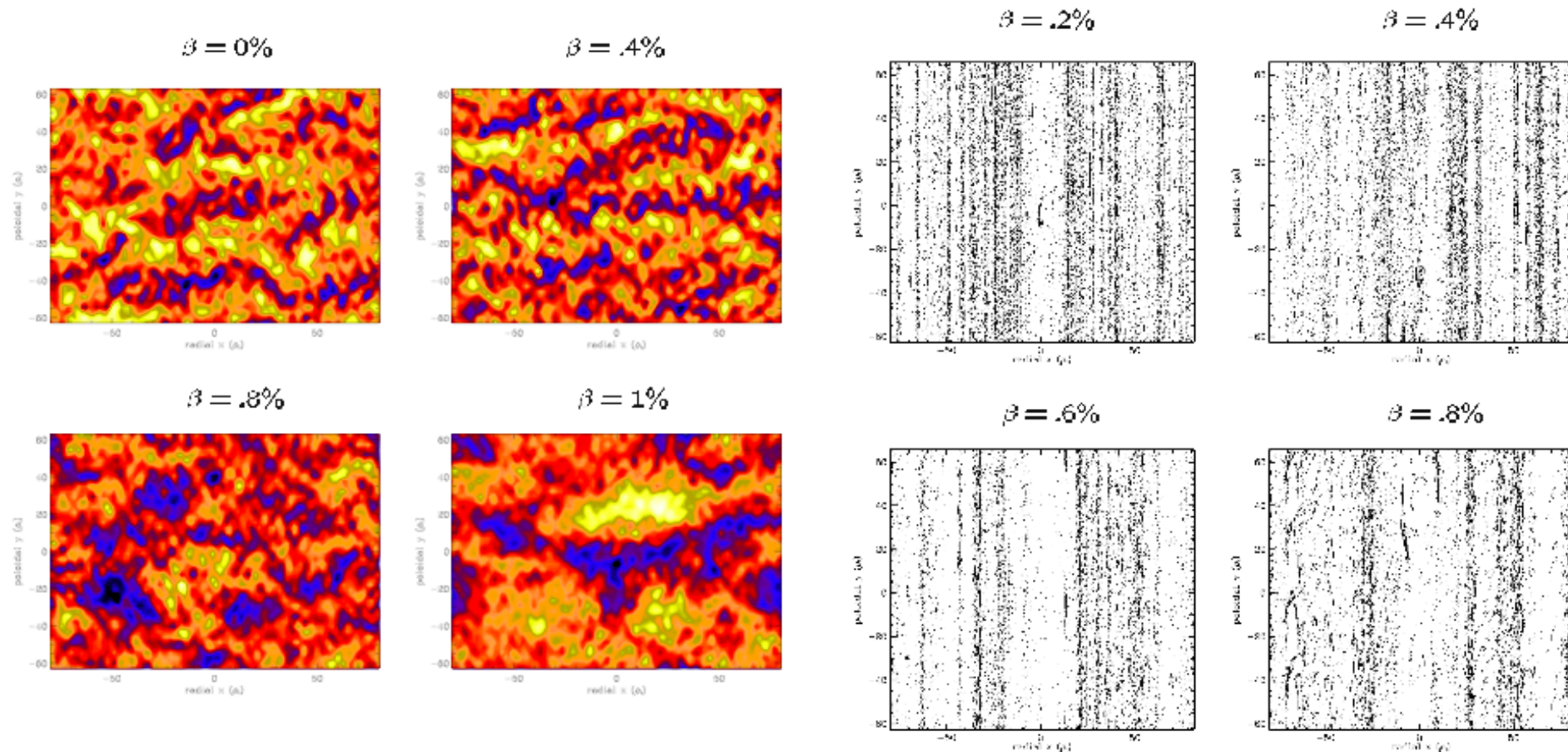
Linear Toroidal ITG Growth Rates  
Comparison with M. Kotschenreuther's GK Code



- Reproduces Growth Rate for finite- $\beta$  ITG as well as kinetic ballooning mode instability
  - KBM unstable below ideal threshold when temperature gradient is finite



# GF Equations used in NL flux tube simulations of EM drift modes



- **Small reduction in flux at low beta, then increase as KBM threshold approached**
  - Character of turbulence changes, B field stochastic, near KBM limit

# Comparison of Methods for Solving GK Eqn

- **GK Eulerian or Continuum methods (eg GS2, GYRO, GENE)**
  - Grid v space and directly solve 5D equation
  - Can choose v space coords and grid for efficiency
  - Collisions straightforward to implement in principle, no noise issue
- **Particle-in-Cell (eg GTC/GTS, Parker's GEM...)**
  - Use markers (sometimes called superparticles or particles) to resolve velocity space in Monte-Carlo-like fashion
  - Relatively straightforward to code and scale, can get good v-space resolution via time averaging
  - Noise issues, and challenging to implement realistic collisions
- **Gyrofluid**
  - Take ~6 moments of GK eqn, kinetic closures, conservative
  - Moments are v-space grid, ~10-100 times more efficient
  - Nonlinear kinetic damping not treated
  - Some closures artificially damp R-H zonal flow, correctable problem
    - Myth: this is source of IFS-PPPL controversy; Reality: relatively small effect

# Issues for GF in edge turbulence and ELM problems

- Small perturbations assumed in closure derivation
- Need extension to full B, kink term
- Much of GF efficiency comes from flux tube, k-space
  - Closure terms become integral operators in real space
    - Efficient evaluation is challenging
    - Simplification (eg localization or Scott constant method)

$$q_{\parallel s} = -8n_0 v_{t\parallel s}^2 \frac{ik_{\parallel} T_{\parallel s}}{(\sqrt{8\pi}|k_{\parallel}|v_{t\parallel s} + (3\pi - 8)\nu_s)}$$

$$q_{\perp s} = -\frac{n_0 v_{t\parallel s}^2 ik_{\parallel} T_{\perp s}}{(\sqrt{\frac{\pi}{2}}|k_{\parallel}|v_{t\parallel s} + \nu_s)} + \left(1 - \frac{T_{\perp 0s}}{T_{\parallel 0s}}\right) \frac{n_0 v_{t\parallel s}^2 T_{\perp 0s} ik_{\parallel} B_1}{B_0(\sqrt{\frac{\pi}{2}}|k_{\parallel}|v_{t\parallel s} + \nu_s)}$$

$$q_{\parallel s}(z) = -n_0 \left(\frac{2}{\pi}\right)^{\frac{3}{2}} v_{t\parallel s} \int_0^{\infty} dz' \frac{T_{\parallel s}(z+z') - T_{\parallel s}(z-z')}{z'},$$

$$q_{\parallel s} = -n_0 \left(\frac{2}{\pi}\right)^{\frac{3}{2}} v_{t\parallel s} \int_0^{\infty} d\hat{z}' g(\hat{z}') [T_{\parallel s}(\hat{z} + \hat{z}') - T_{\parallel s}(\hat{z} - \hat{z}')] ]$$

$$g(\hat{z}) = \int_0^{\infty} d\hat{k} \frac{\hat{k}}{\hat{k} + 1} \sin(\hat{k}\hat{z})$$



# Discussion

- **Principal advantage of GF eqs is their efficiency and ability to incorporate collisionless damping**
  - Also relatively easy to work with and simplify in various limits
  - Generally well behaved numerically, conservative
  - Easier than GK to interpret results
  - Right compromise between accuracy and efficiency?
- **Weakness relative to direct GK is additional simplifications**
  - No nonlinear Landau damping, drift res not exact
  - One model is GF/GK working in tandem for efficiency
- **Weakness relative to Braginskii-like eqs is presence of closure terms which are non-local in real space**
  - Braginskii assumes very high collisionality (and can be poorly behaved at low collisionality). GF assumes GK ordering is valid.

- **Extra slides**

# Some generalization required for MHD-like problems

- Need to avoid simplified equilibrium, keep full B perturbation and kink term

The gyrocenter velocity is then given by

$$\dot{\mathbf{X}} = v_{\parallel}(\hat{\mathbf{b}} + \frac{\langle \delta \mathbf{B}_{\perp} \rangle}{B}) + \mathbf{v}_E + \mathbf{v}_d, \quad (3.3)$$

where the angular brackets denote gyroangle averages. The first term on the right represents free streaming along the total magnetic field. The second term is the gyroaveraged  $\mathbf{E} \times \mathbf{B}$  drift velocity,  $\mathbf{v}_E = \frac{c}{B} \hat{\mathbf{b}} \times \nabla \langle \phi \rangle$ .  $\mathbf{v}_d$  is the combined curvature and  $\nabla B$  drift velocity. In general,  $\mathbf{v}_d$  can be written

$$\begin{aligned} \mathbf{v}_d &= \frac{v_{\parallel}^2}{\Omega} \hat{\mathbf{b}} \times (\hat{\mathbf{b}} \cdot \nabla \hat{\mathbf{b}}) + \frac{\mu}{\Omega} \hat{\mathbf{b}} \times \nabla B \\ &= \frac{v_{\parallel}^2 + \mu B}{\Omega B^2} \mathbf{B} \times \nabla B + \frac{v_{\parallel}^2}{\Omega B^2} \hat{\mathbf{b}} \times (\nabla \times \mathbf{B} \times \mathbf{B}). \end{aligned} \quad (3.4)$$

Using the equilibrium relations  $\nabla p = \frac{1}{c} \mathbf{J} \times \mathbf{B}$  and  $\nabla \times \mathbf{B} = \frac{4\pi}{c} \mathbf{J}$ , this can be written

$$\mathbf{v}_d = \frac{v_{\parallel}^2 + \mu B}{\Omega B^2} \mathbf{B} \times \nabla B + \frac{v_{\parallel}^2}{\Omega B^2} \hat{\mathbf{b}} \times \nabla p. \quad (3.5)$$

The second term on the right is small for  $\beta \ll 1$ ,<sup>3</sup> and is neglected here for simplicity and to maintain consistency with neglecting  $\delta B_{\parallel}$ .<sup>4</sup> The definition

$$\mathbf{v}_d \doteq \frac{v_{\parallel}^2 + \mu B}{\Omega B^2} \mathbf{B} \times \nabla B \quad (3.6)$$